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# Unmasking the Mysteries of High-Mass X-Ray Binaries (HMXBs): The Role of LLNL's Electron Beam Ion Trap (EBIT)

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## GOAL

- To analyze X-ray line spectra of excited, highly charged silicon (Si) produced by EBIT.

## PURPOSE

- To use data as a point of reference for similar spectra measured by satellites *Chandra* and *XMM-Newton*.
- To calculate Doppler shift of X-ray lines and, in turn, ion velocity around HMXBs.

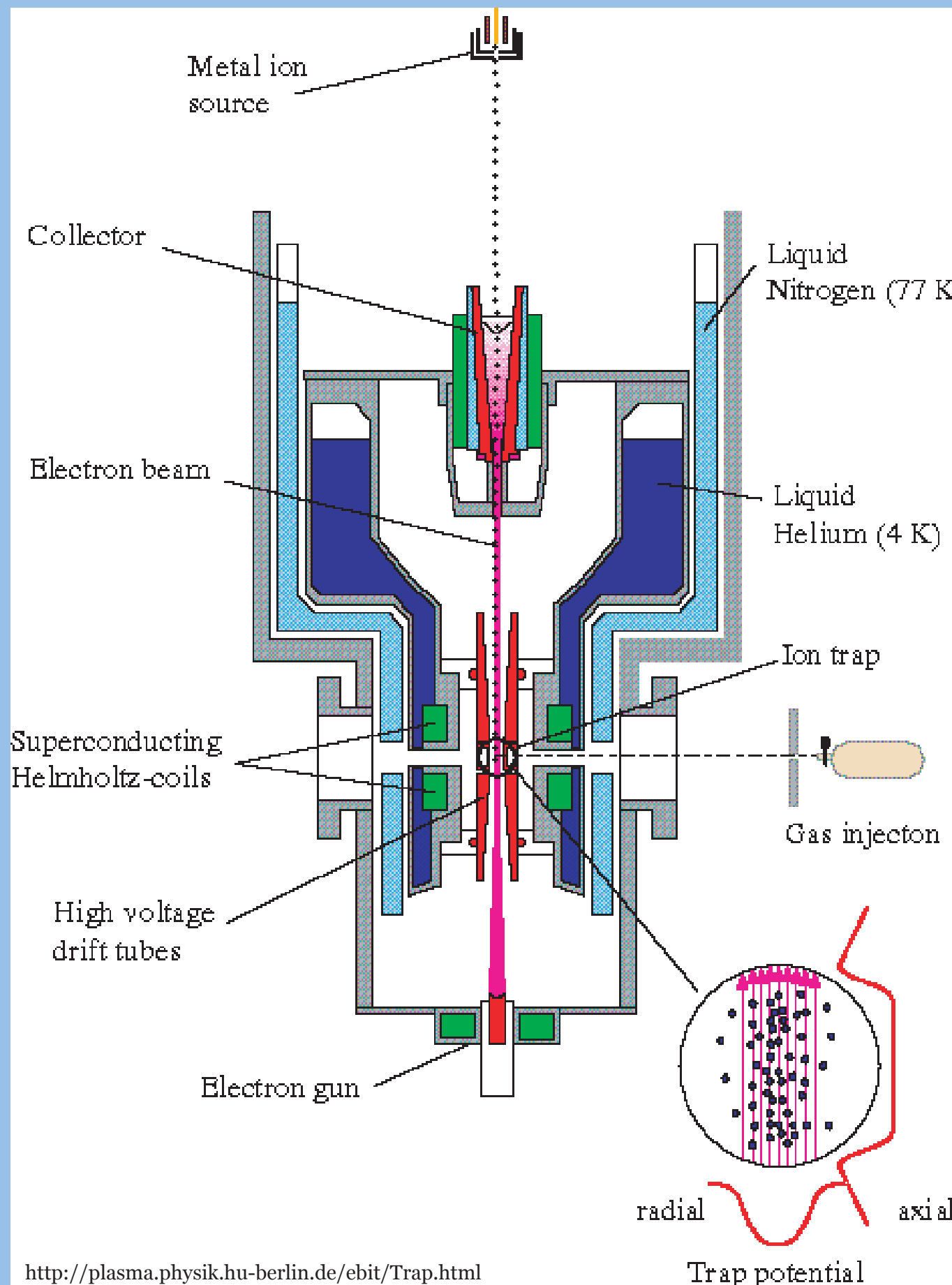
## BACKGROUND

### EBIT

- EBIT uses a narrow electron (~60μm) beam to excite and trap ions.
- X-ray emission from the excited ions is then diffracted using crystal spectrometers and analyzed.

### WHY EBIT IS IMPORTANT

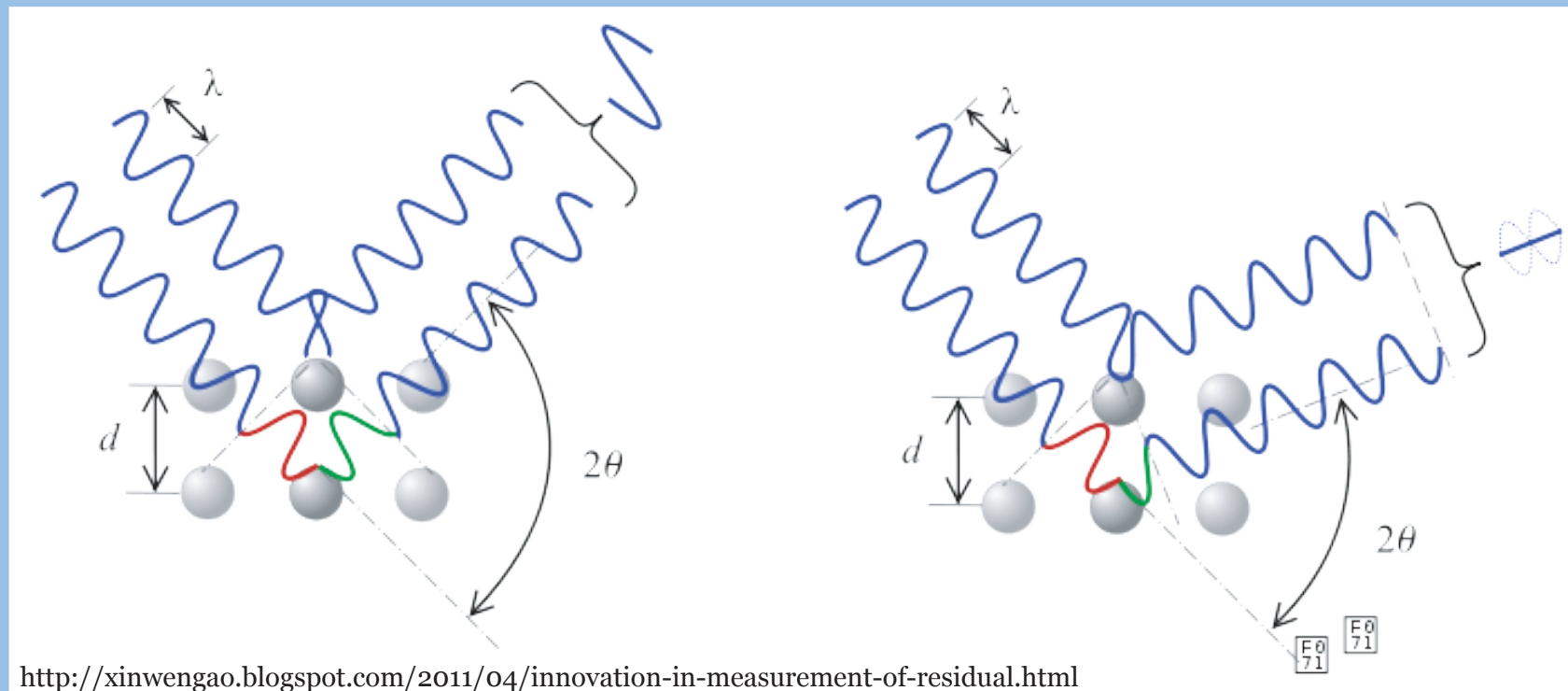
- High precision, high accuracy
- Provides spectral data that can't be calculated
- Helps astronomers to better understand wind movement around HMXB accretion disks



## CRYSTAL SPECTROMETERS DIFFRACT X-RAYS ACCORDING TO BRAGG'S LAW

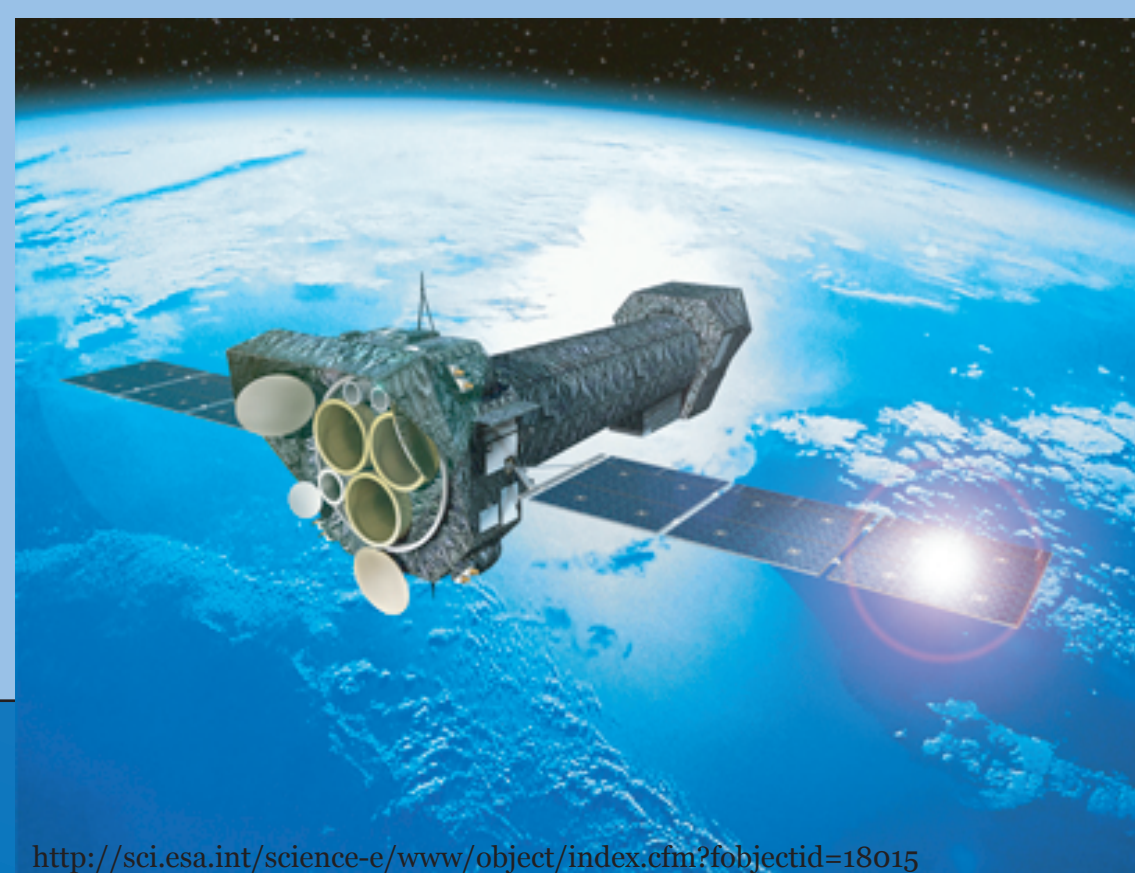
$$N\lambda = 2d\sin\theta$$

N – integer  
λ – wavelength of incident wave  
d – space between lines or atoms in target material  
θ – angle between incident ray and scattering planes



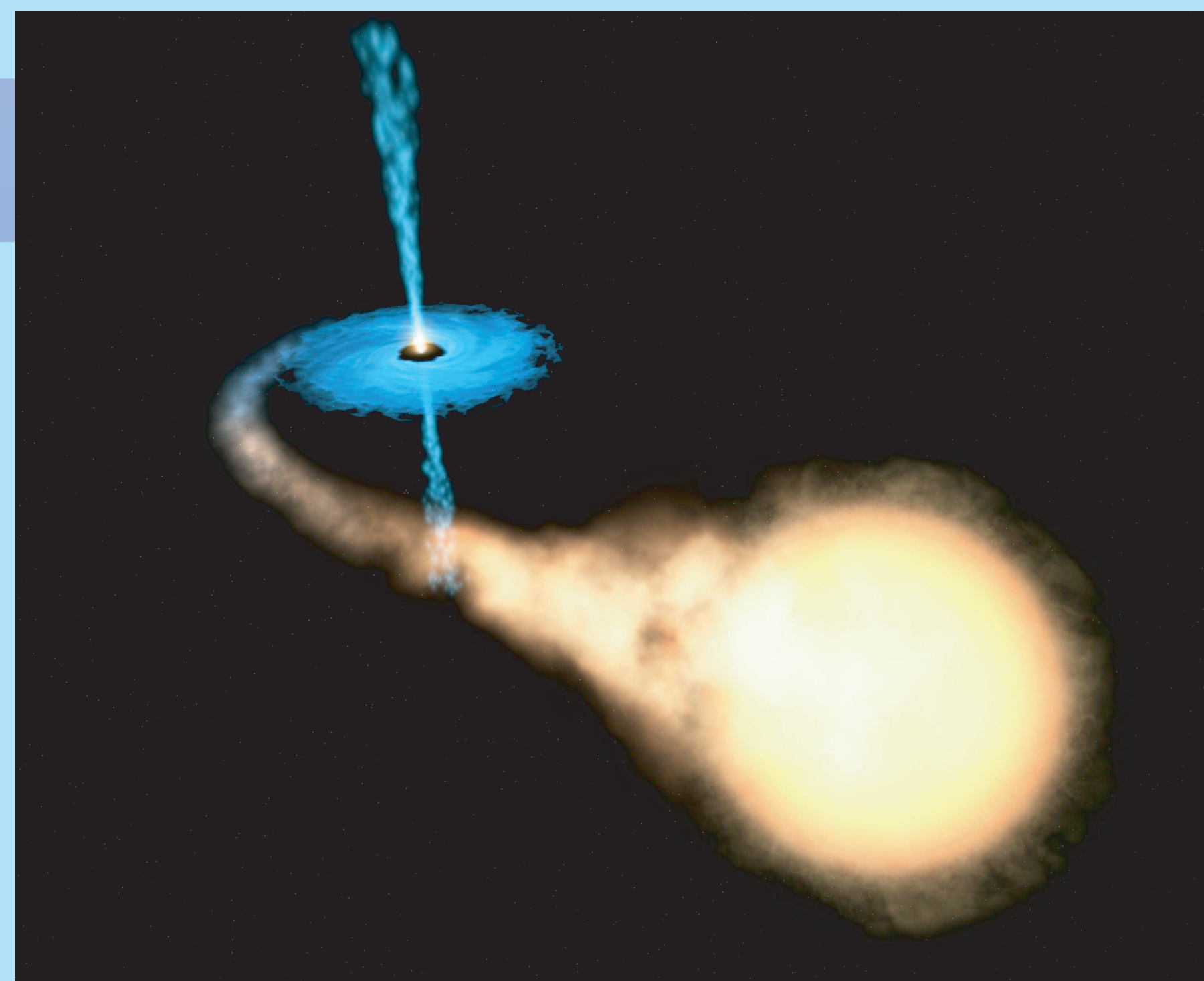
Bragg diffraction occurs when electromagnetic waves (like X-rays) encounter obstacles (like atoms in a crystal) whose spacing is comparable to the wavelength.

Satellites  
*Chandra*  
(left) and  
*XMM-Newton*.



## APPLICATION

- Monoenergetic electron beam allows for isolation of single charge state.
- Can probe different excitation mechanisms.
- Accuracy of better than ~5mÅ is achievable.
- Can calculate Doppler shifts of Si spectra and learn about wind around HMXBs' accretion disks.

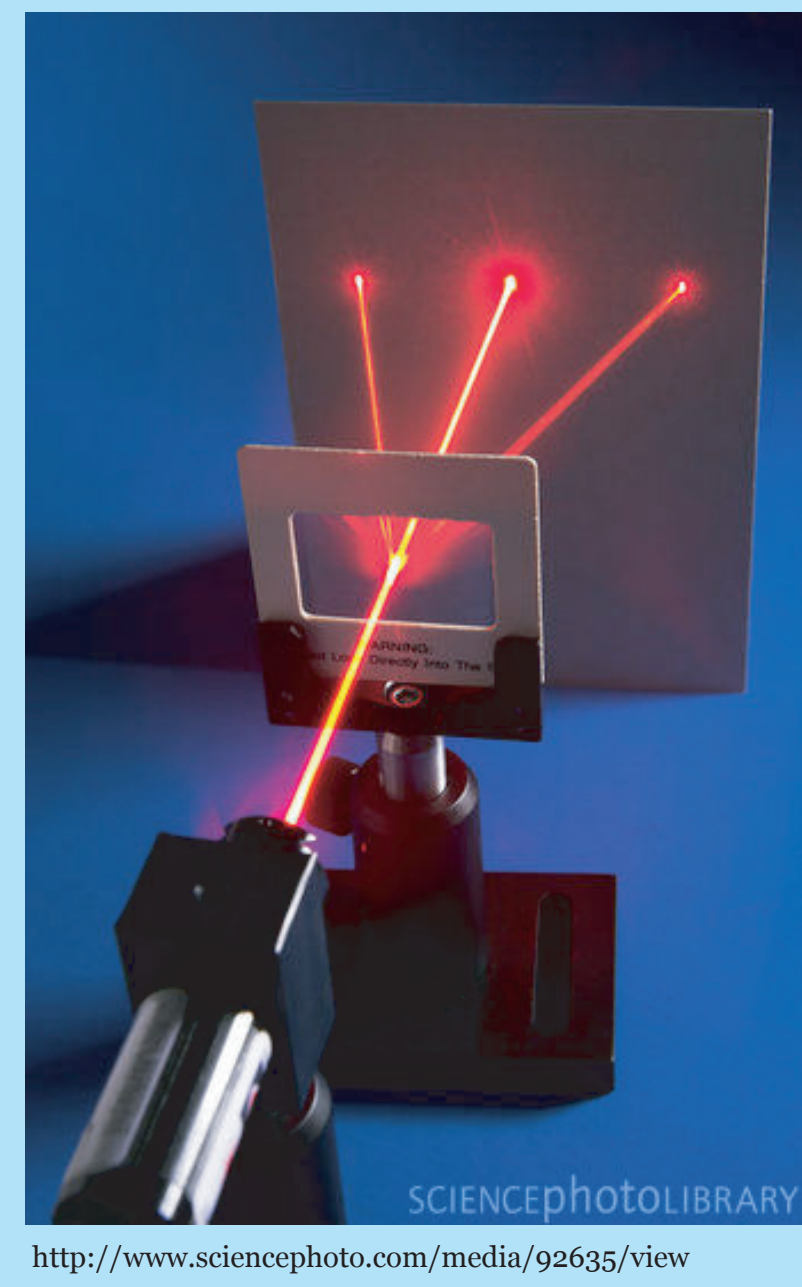


Artist's rendering of a HMXB: A compact object (left, neutron star or black hole) and its companion, or donor, star.

## IN THE CLASSROOM

### The Nature of Light: Diffraction and Interference

Can you explain this?



- Diffraction* – the bending of waves around an object
- Interference* – a phenomenon in which two waves superimpose to form a resultant wave of greater or lower amplitude

- Challenge: Use these concepts to:

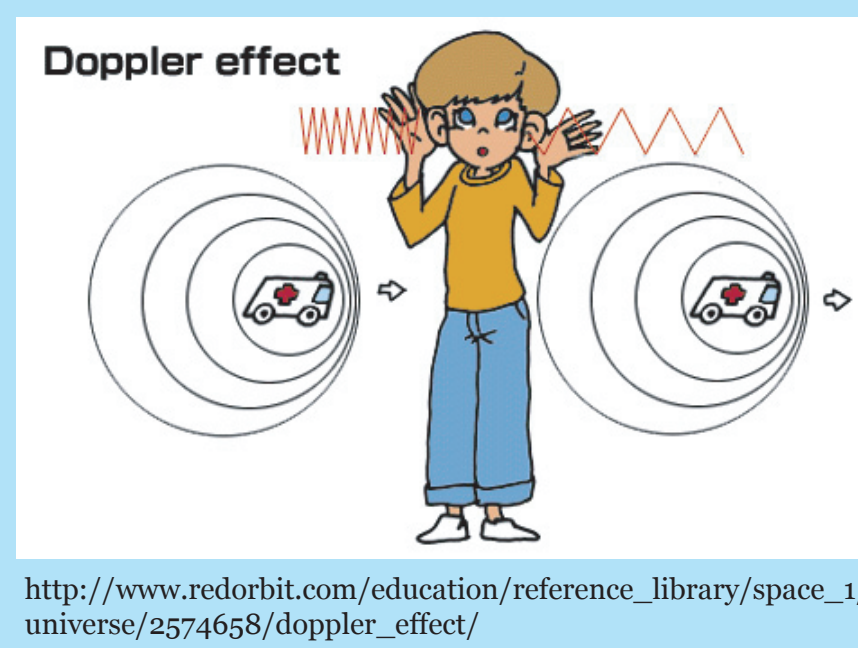
- Find track spacing of CD and DVD
- Find diameter of lycopodium spore

### Wave Behavior: The Doppler Shift

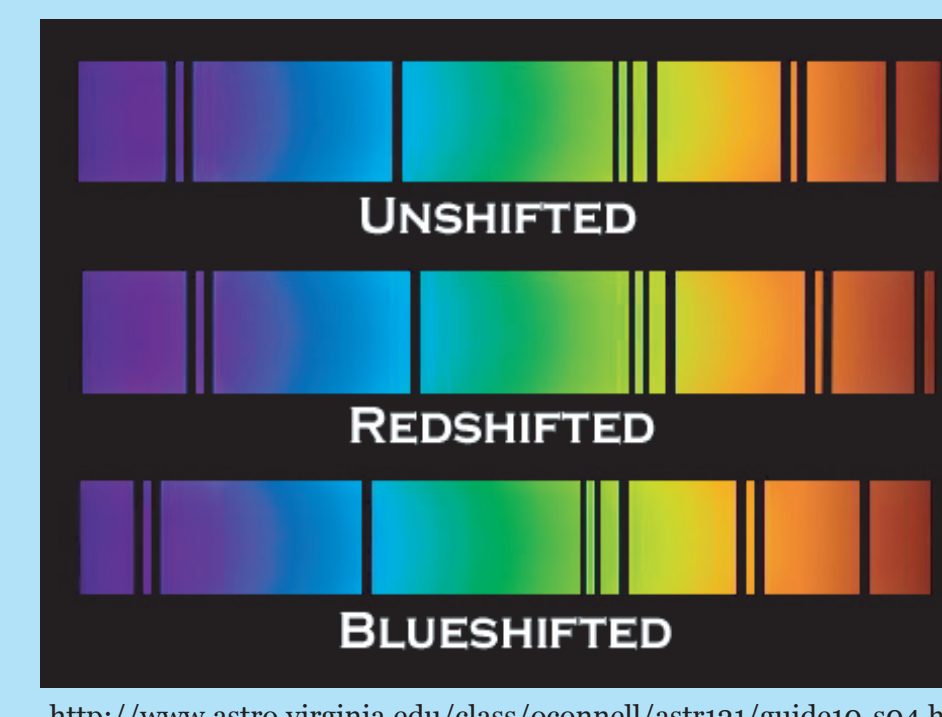
$$\frac{\Delta\lambda}{\lambda_0} = \frac{v}{c}$$

Δλ = wavelength shift  
λ<sub>0</sub> = wavelength of source not moving  
v = velocity of source  
c = speed of light

Sound waves

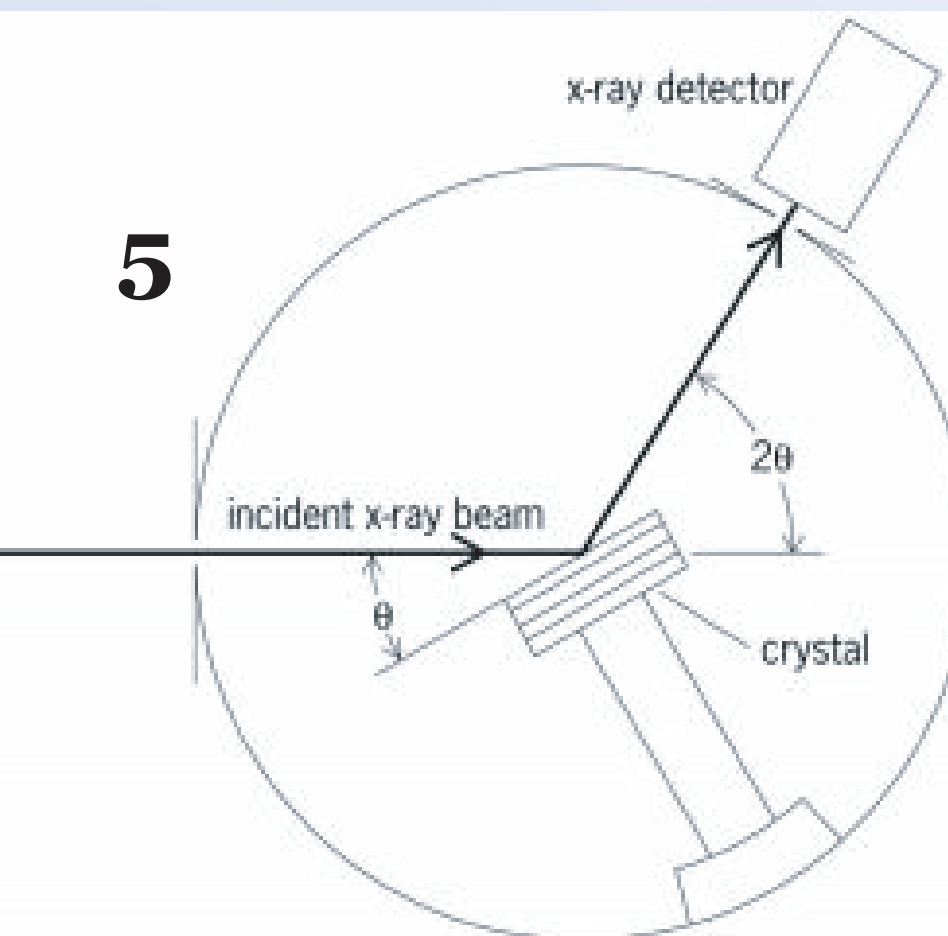


Light waves



## PROCEDURE

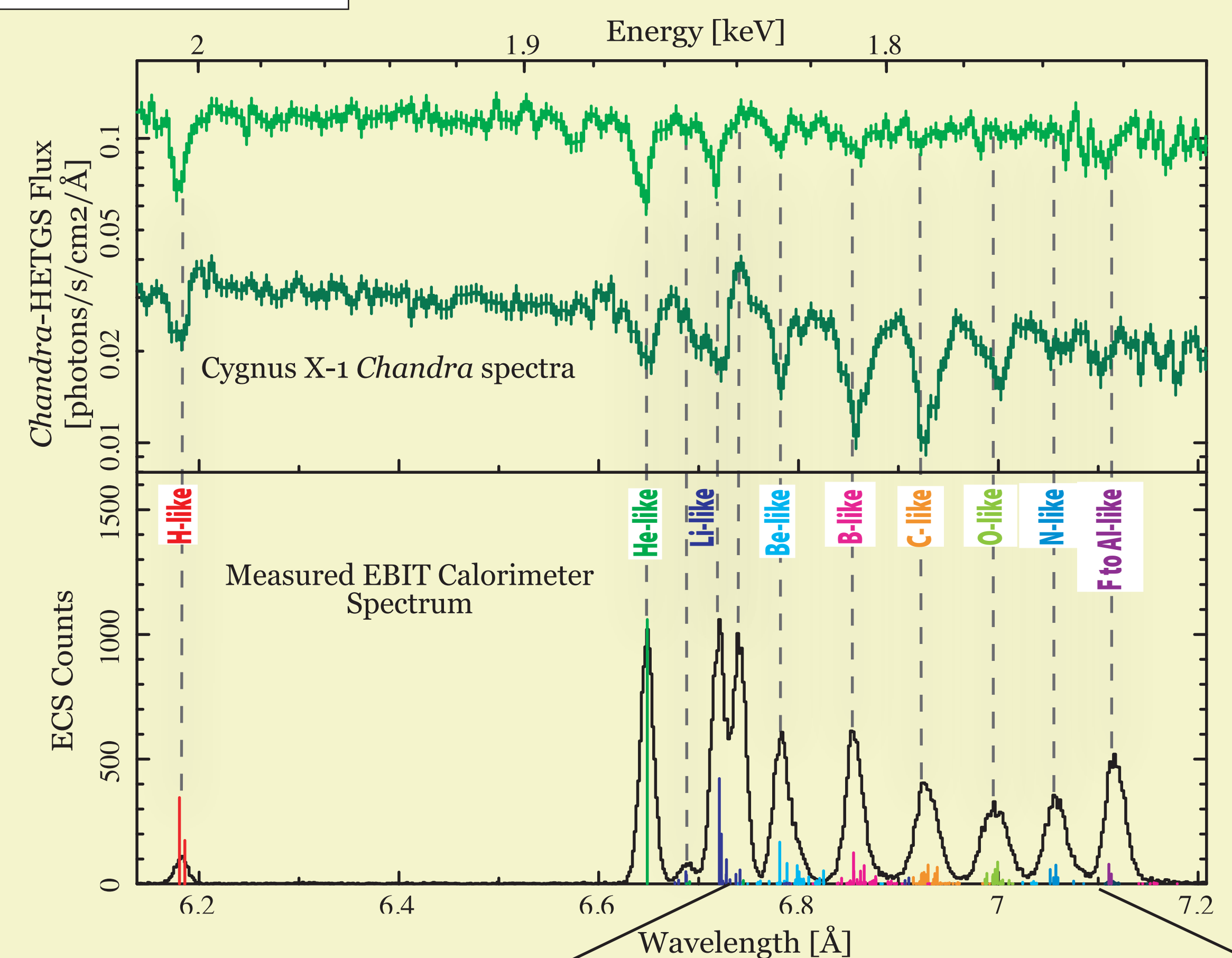
- Injected Si (in the form of *decamethyltetrasiloxane*) into EBIT.
- Used EBIT's monoenergetic electron beam to produce and excite highly charged Si ions Si<sup>5+</sup> to Si<sup>12+</sup>.



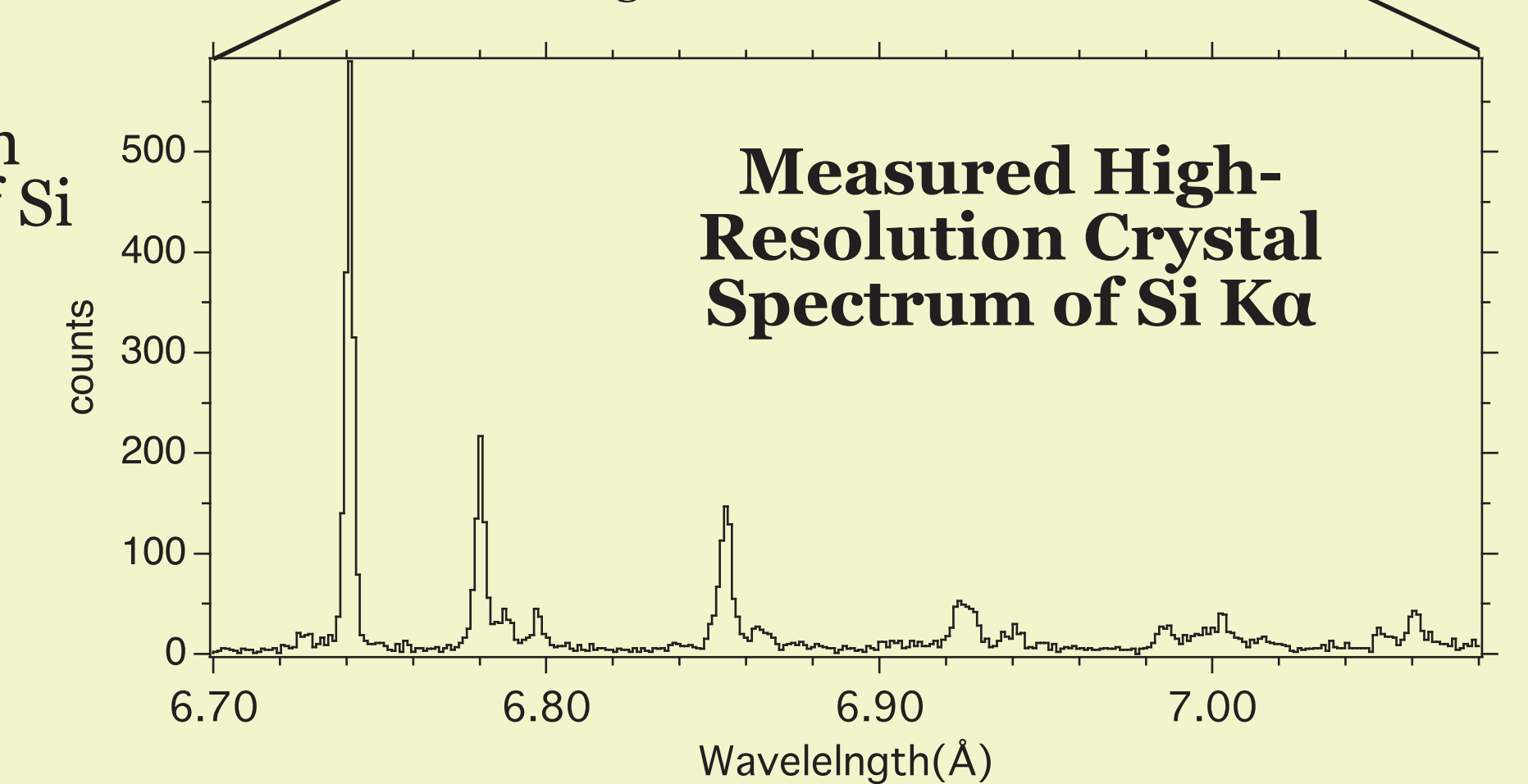
- The strongest 2 to 1 transitions in each charge state were measured using a crystal spectrometer and an X-ray calorimeter.
- The wavelength scale was calibrated using well-known wavelengths from hydrogenic and helium-like ions.

## RESULTS

Si Kα Spectra: EBIT-Generated Emission compared to Absorption in Cygnus X-1 Observed by *Chandra*



The high-resolution crystal spectrum of Si resolves blends in Calorimeter and *Chandra* spectra.



- Absorption lines of highly charged Si ions have been observed in the spectra of Cygnus X-1.
- In order to determine if the measured absorption lines are Doppler-shifted, rest energy line centers are needed.
- The lower panel shows the Kα transitions of several charge states of Si measured with the EBIT Calorimeter Spectrometer (ECS) at LLNL while the ions are at rest.
- Overlaid are the color-coded theoretical calculations of these lines.
- Some Si ions are Doppler shifted. Ion velocities of ~200 km/s were determined.

<http://www.nasa.gov/vision/universe/stargalaxies/chandra2years.html>

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