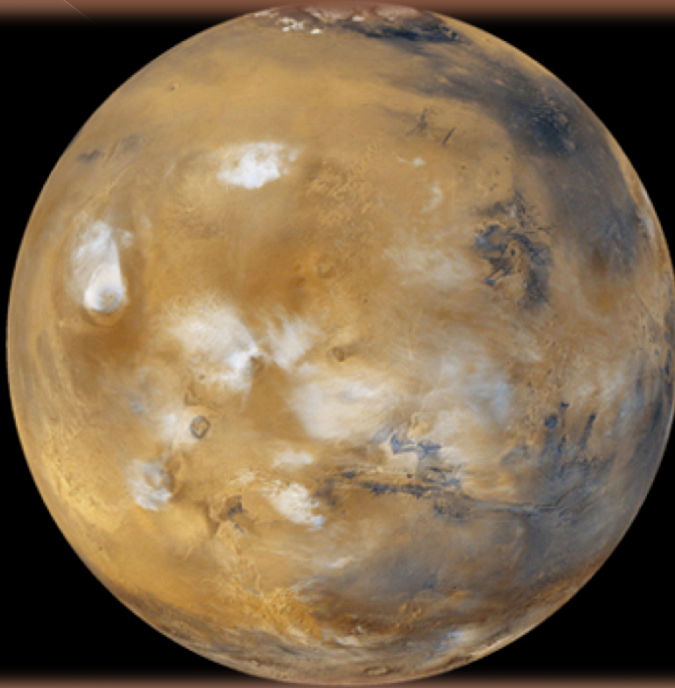


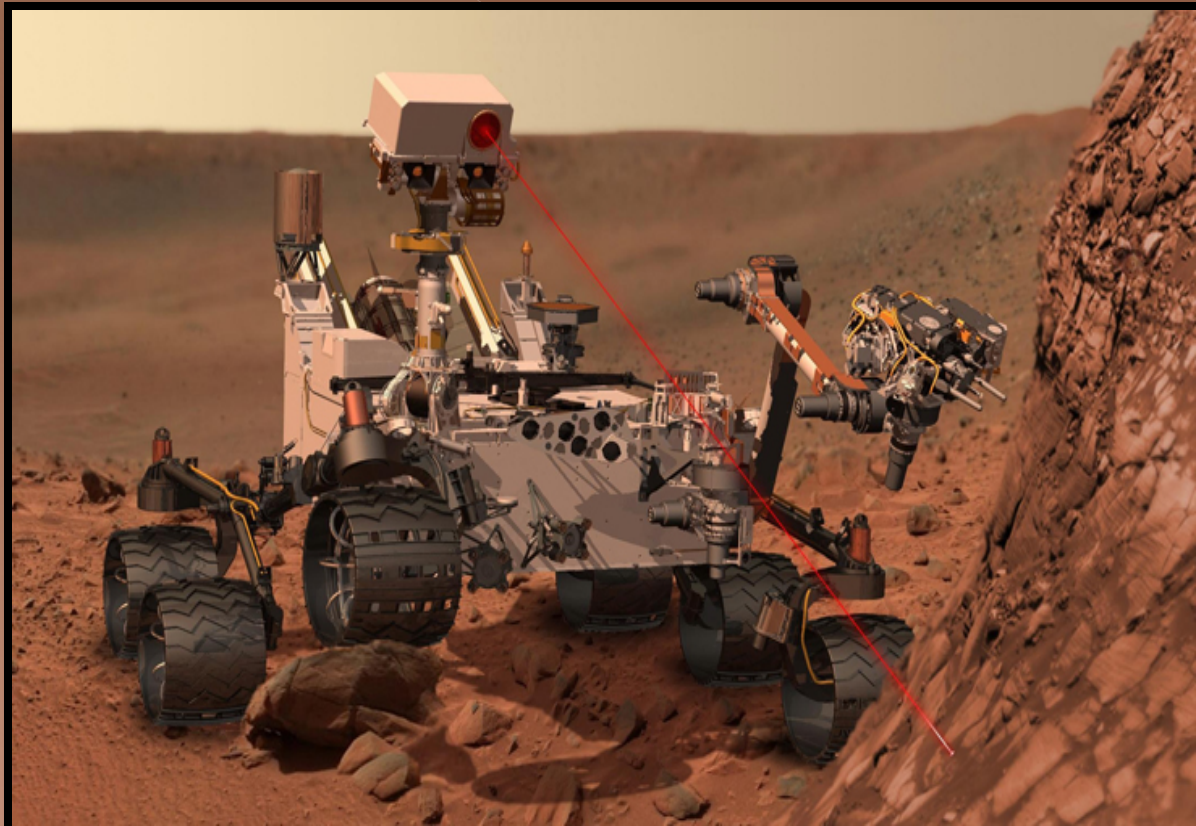
# *Raman Mapping of Mars Relevant Minerals Using Multiple Excitation Wavelengths*



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Jet Propulsion Laboratory,  
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Mentor: William Abbey  
Summer 2012  
CSU STAR Program

# *The Ultimate Question*

- How do we define the next generation in-situ instrument used to map minerals on Mars?



# *Why does this matter?*

## *“Big Picture”*

This study is meant to eventually contribute to...

1. Non-invasive ‘hands’ free sampling
2. Detection and spatial correlation between minerals and trace organics
3. Spatial relationships which may provide contextual information about the origin of detected materials

# *Methodology*

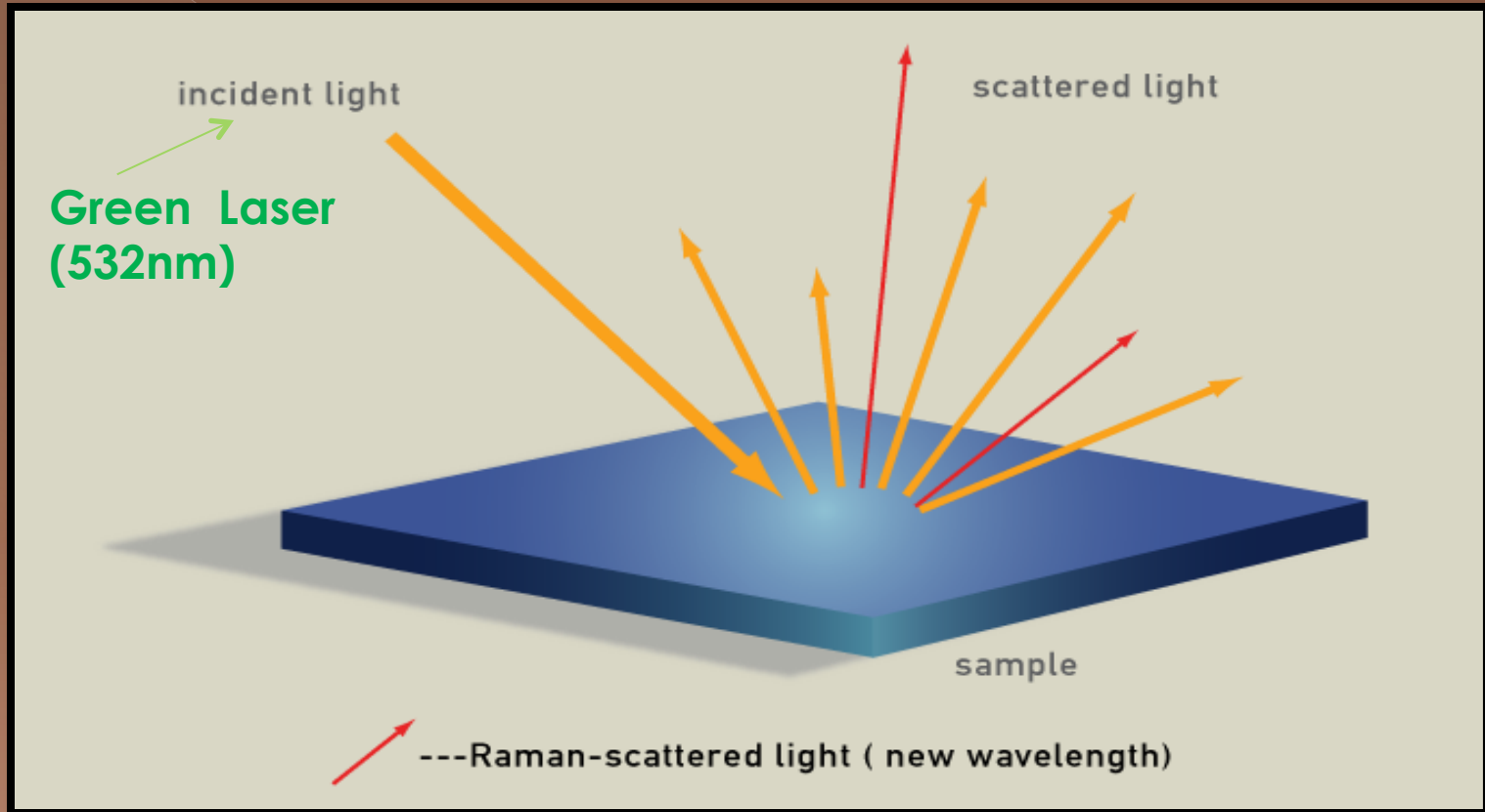
- Read about Raman Spectrometry and research Mars relevant minerals
- Create powdered samples for analysis of composition
- Use Raman Spectrometer given specific parameters (spot size, mineral types)
- Create data base for minerals



# *Why Raman Spectroscopy?*

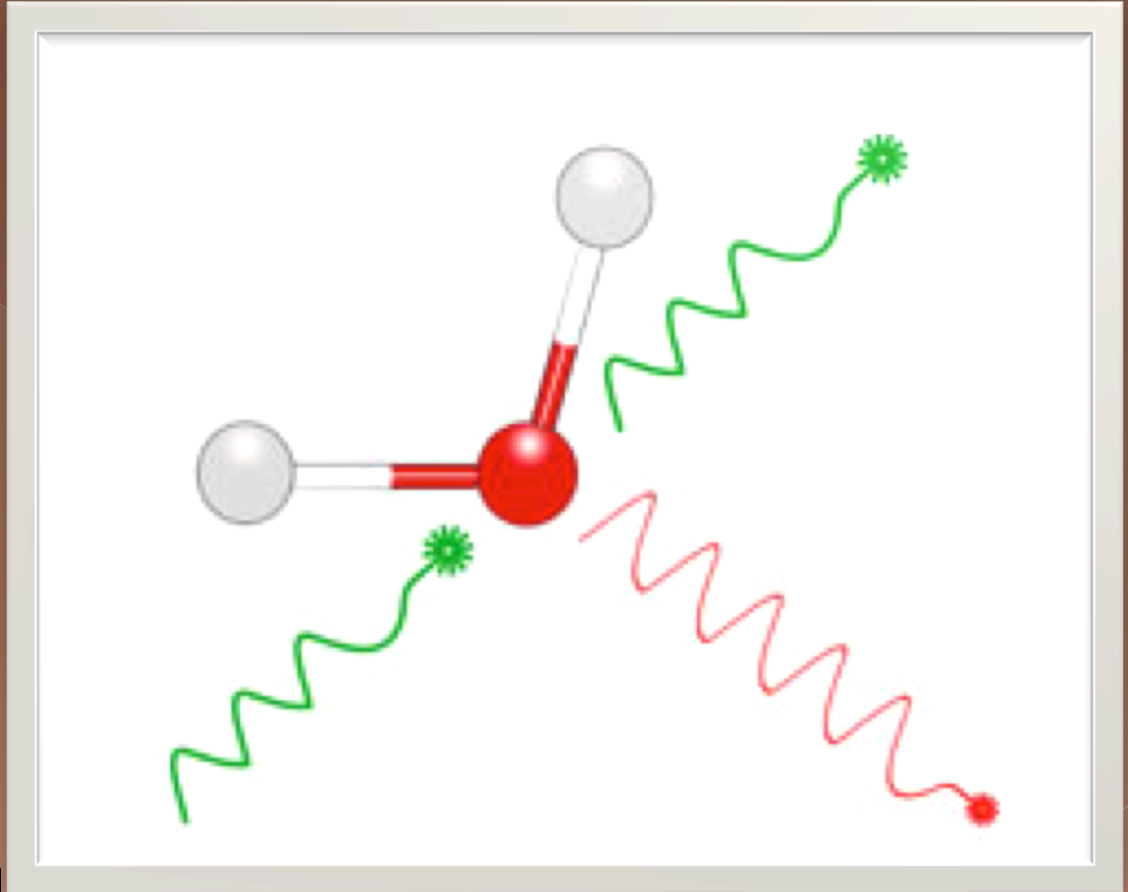
- The Raman spectrometer is a non-invasive sampling technique used to determine the 'Raman Shift' of each mineral
- The system provides high spectroscopic resolution and a unique wavelength range capability that offers both great flexibility and high performance.

# How the Raman works



# Light scattered from a molecule

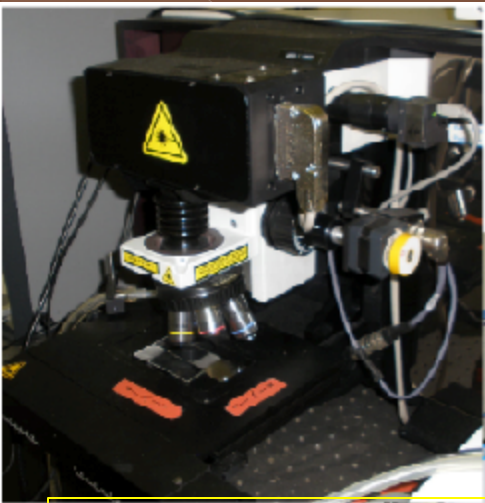
The incident photons interact with the molecules, and the amount of energy change (either lost or gained) by a photon is characteristic of the nature of each bond vibration present.



Green \_\_\_\_\_  
Incident Rays

Red \_\_\_\_\_  
Scattered Rays

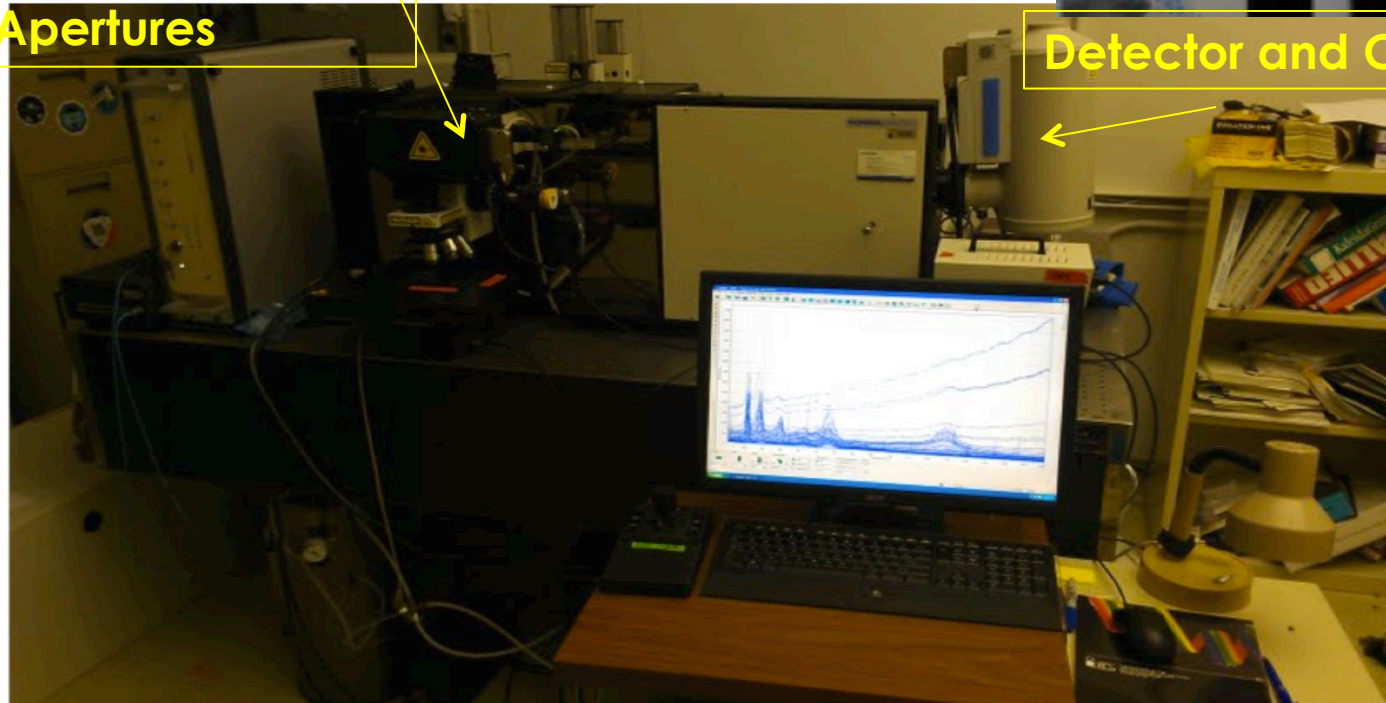
# *The Spectrometer*



Camera, Lenses and  
Laser Apertures



Detector and Chiller



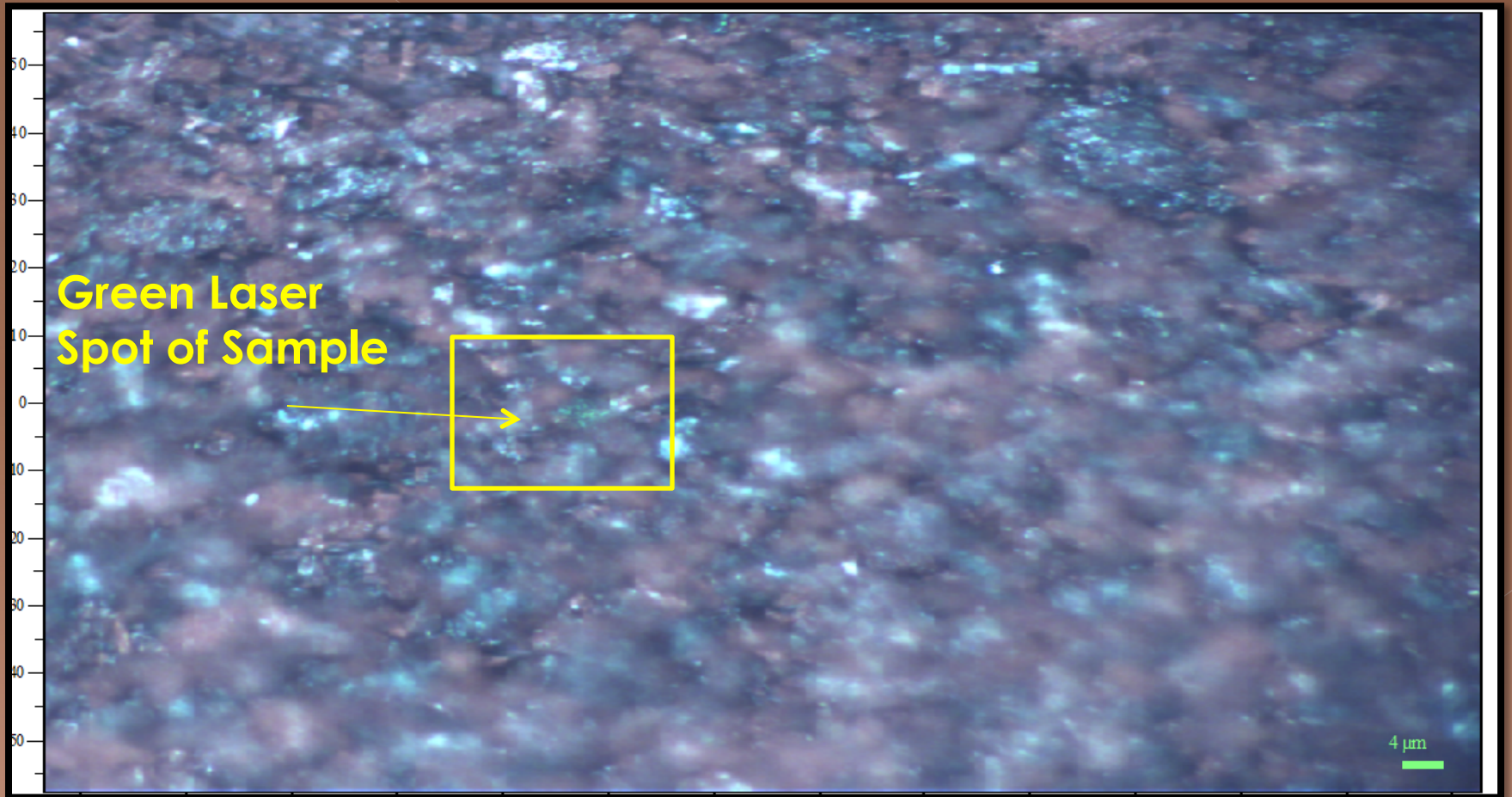


# *List of Minerals Tested*

- ◉ **Enstatite**  $\text{MgSiO}_3$  **Mg** Silicate
- ◉ **Forsterite**  $\text{MgSiO}_4$  **Mg** Silicate
- ◉ **Hematite**  $\text{Fe}_2\text{O}_3$  **Iron** Oxide
- ◉ **Magnetite**  $\text{Fe}_3\text{O}_4$  **Iron** Oxide
- ◉ **Jarosite**  $\text{KFe}^{3+}_3(\text{OH})_6(\text{SO}_4)_2$  Hydrous Sulfate
- ◉ **Gypsum**  $\text{CaSO}_4 \cdot 2\text{H}_2\text{O}$  Hydrous Sulfate
- ◉ **Halite**  $\text{NaCl}$  Salt
- ◉ **Calcite**  $\text{CaCO}_3$  Carbonate
- ◉ **Pyrite**  $\text{FeS}_2$  **Iron** Sulfide

# *Hematite Powder*

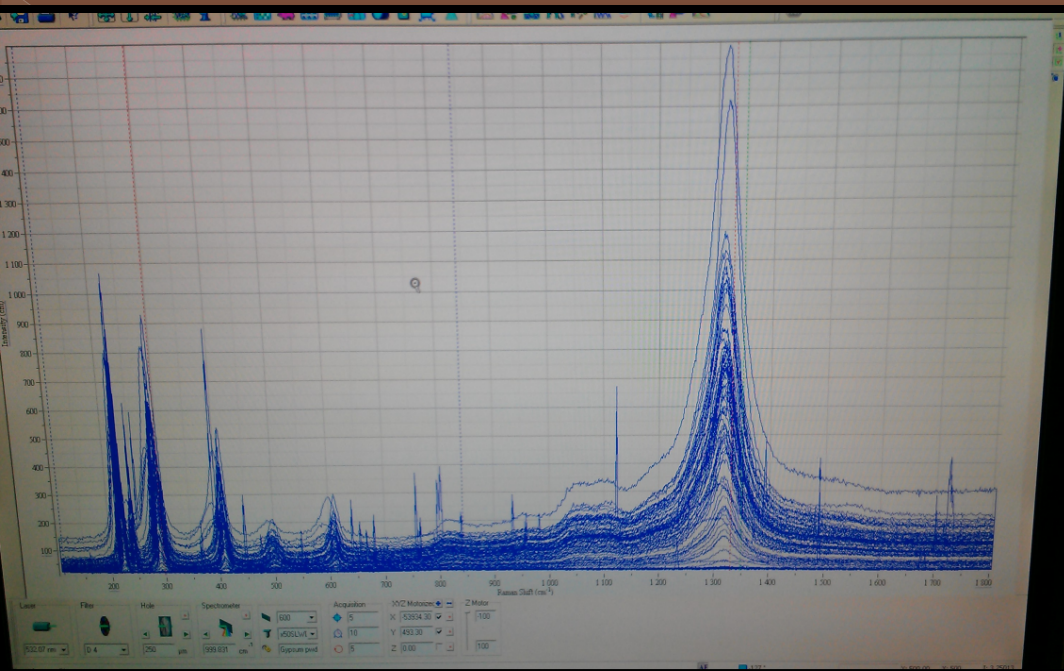
## *$\text{Fe}_2\text{O}_3$ Iron Oxide*



# ***Mars Relevant Minerals***

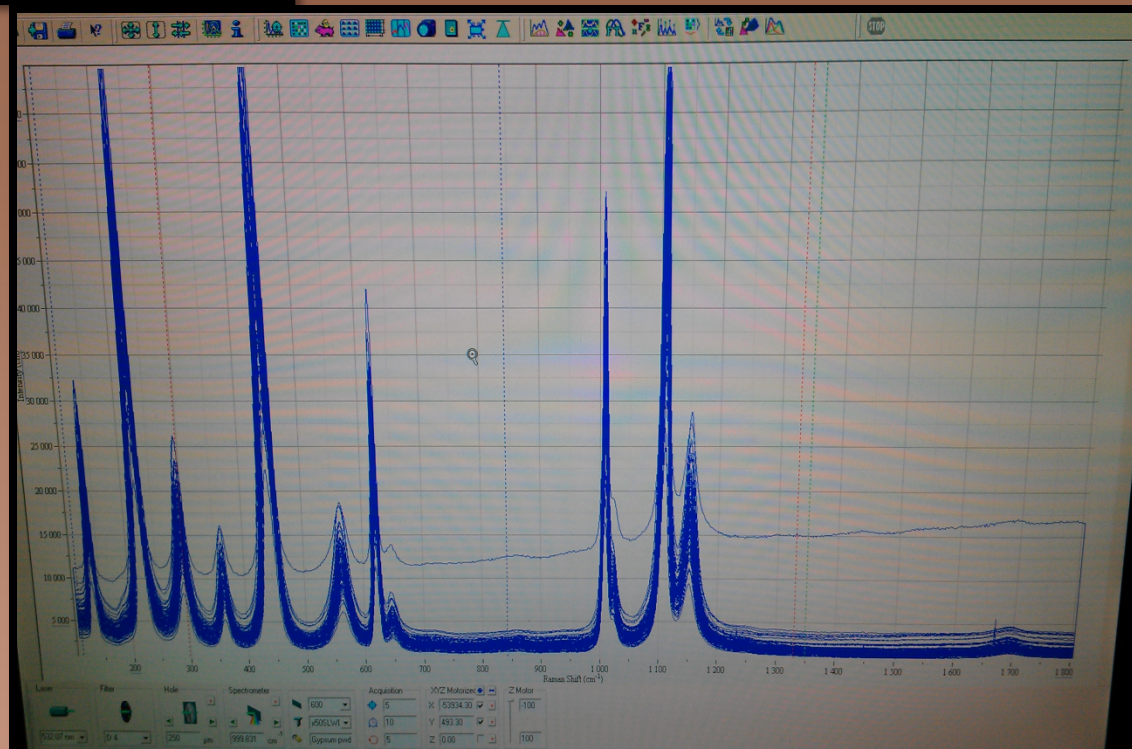
- Each mineral has a unique spectrum with peaks at certain wavelengths.
- By examining the spectrum and making note of where significant peaks are located, we can identify what mineral is present.





Hematite  
 $(\text{Fe}_2\text{O}_3)$   
 Iron Oxide

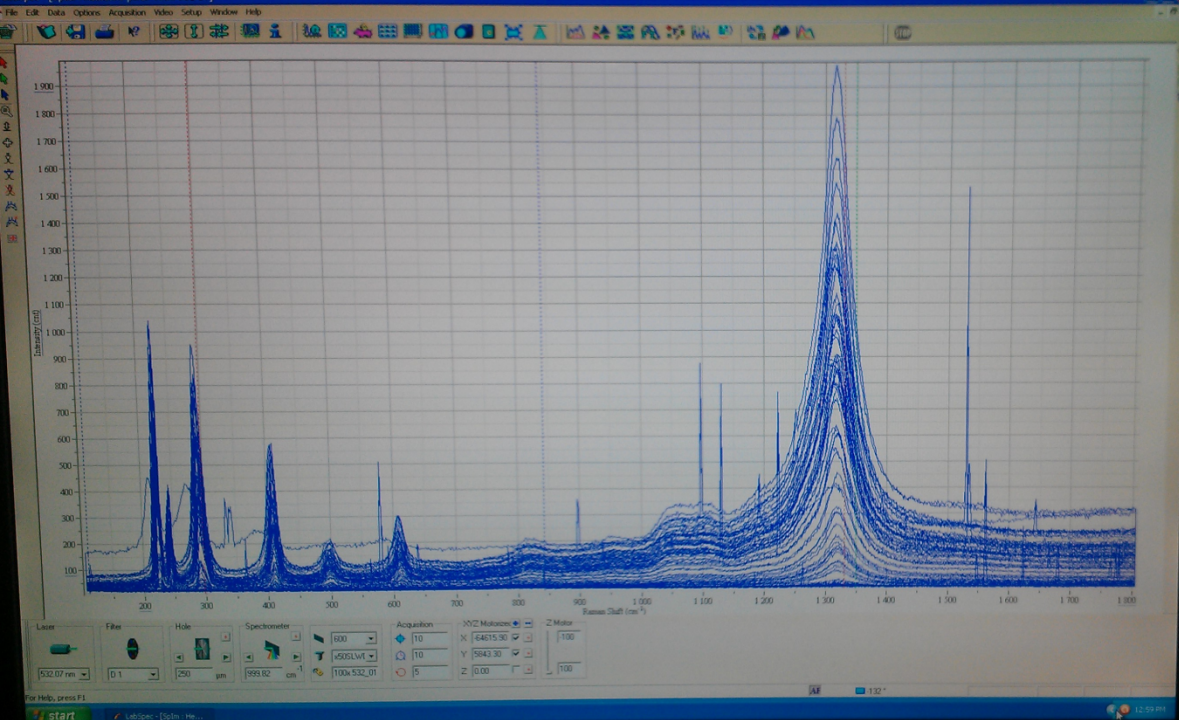
NatroJarosite  
 $\text{KFe}^{3+}_3(\text{OH})_6(\text{SO}_4)_2$   
 Hydrated Sulfate





# *Challenges*

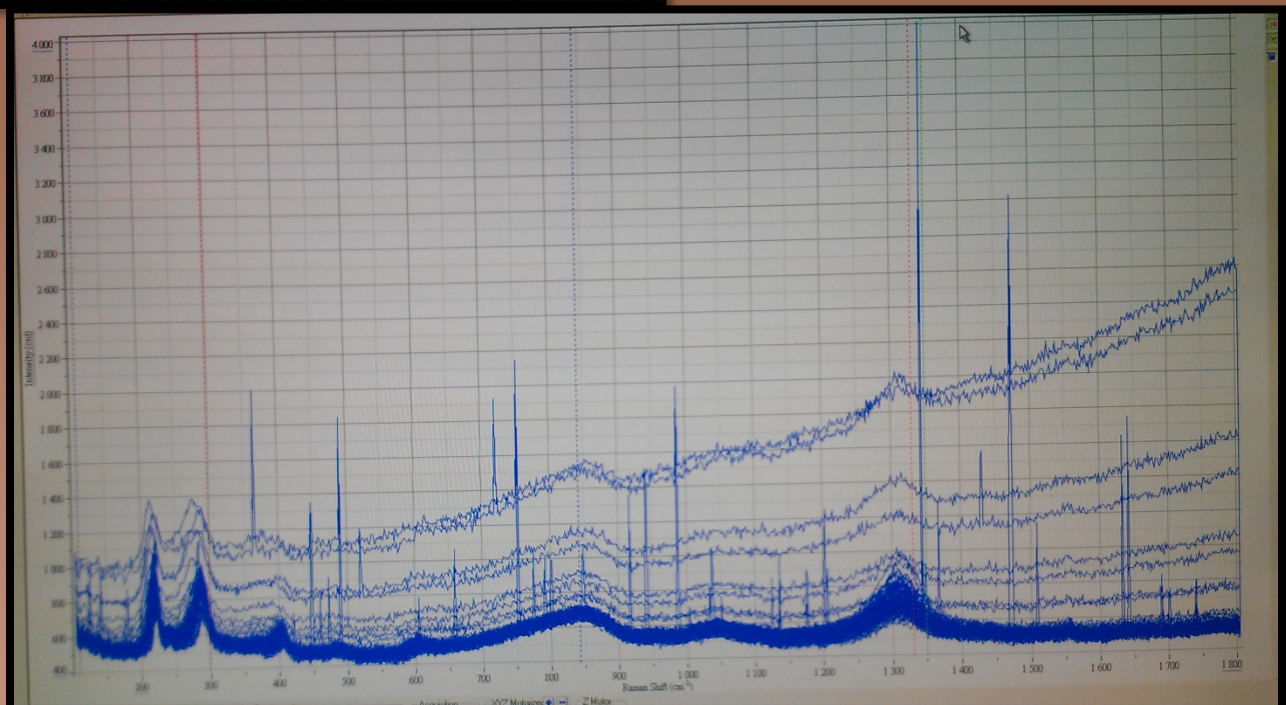
- ◉ Raman calibration
- ◉ Preparing the samples properly
- ◉ Grain alignment (clays)
- ◉ Time it takes to run scans and to clean the samples
- ◉ Obtaining data for 50x and 100x magnification
- ◉ Solar Flares



Hematite  
100x

Variation within a sample  
due to lens size:  
10x vs. 100x

Hematite  
10x



# *Continued Research*

- Switch to Red Laser(632 nm) then UV laser (325 nm)
- Compare data for each laser
- Determine which laser will work best for certain minerals



# *Predicted outcomes*

- Looking for patterns in groups of minerals  
ex: Mg vs. Fe
- Identify which wavelength is best for different suites of minerals
- Use these parameters to determine the type of spectrometer used on next mission

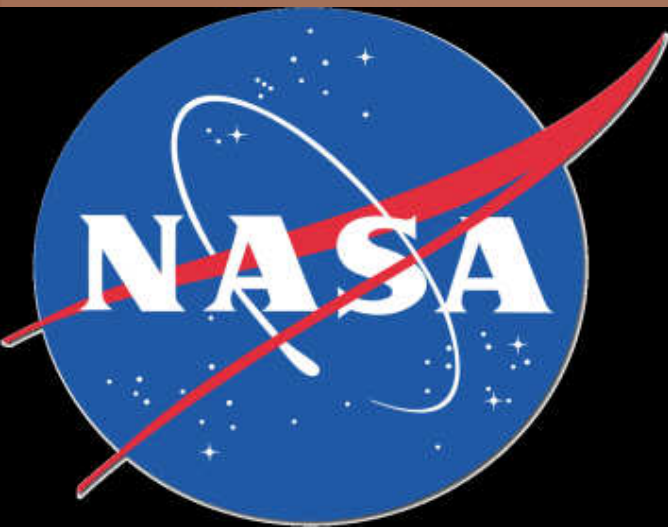


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